# ON THE MOTION abOUT A FIXED POINT OF A fast Spinning heavy solin 

## (0 dvizhenil Privedennogo $v$ bystroe vrashchenie tiAZHELOGO TVERDOGO TELA VOKRUG NEPODVIZHNOI TOCHKI)

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The motion of heavy solid about a fixed point with high initial angular velocity has been investigated (omitting the case of Lagrange) for the case of Goriachev-Chaplygin in [1,2]. Cases with restrictions on the location of the center of gravity, on moments of inertia and on the initial conditions has been investigated in $[3,4]$.

In our work we apply the method of small parameters to investigate the periodic solutions of the equations of motion of a heavy solid with one point fixed rapidly spinning about one of the principal axes of the ellipsoid of inertia. In particular, we show that, with the exception of special cases, when $z_{0} \neq 0$ the body will perform a pseudo-regular precession about the vertical axis in the first approximation, and that at least four of the six initial conditions are arbitrary.

1. Let us consider a heavy solid with one point fixed, whose ellipsoid of inertia is arbitrary, and whose center of gravity is arbitrarily located and not necessarily coinciding with the fixed point. General equations of motion of this solid and their first integrals are

$$
\left.\begin{array}{c}
\frac{d p}{d t}+A_{1} q r=\frac{M g}{A}\left(y_{0} \gamma^{\prime \prime}-z_{0} \gamma^{\prime}\right), \quad \frac{d \gamma}{d t}=r \gamma^{\prime}-q \gamma^{\prime \prime} \quad\binom{A B C, p q r}{\gamma \gamma^{\prime} \gamma^{\prime}, x_{0} y_{0} z_{0}}  \tag{1.1}\\
\left(A_{1}=\frac{C-B}{A}, \quad B_{1}=\frac{A-C}{B}, C_{1}=\frac{B-A}{C}\right) \\
A p^{2}+B q^{2}+C r^{2}-2
\end{array}\right)
$$

$$
\begin{gather*}
A p \gamma+B q \gamma^{\prime}+C r \gamma^{\prime \prime}=A p_{0} \gamma_{0}+B q_{0} \gamma_{0}^{\prime}+C r_{0} \gamma_{0}^{\prime \prime} \\
\gamma^{2}+\gamma^{\prime 2}+\gamma^{\prime 2}=1 \tag{1.2}
\end{gather*}
$$

Here $p_{0}, q_{0}, r_{0}, \gamma_{0}, \gamma_{0}{ }^{\prime}$ and $\gamma_{0}{ }^{\prime \prime}$ are the intial values of the corresponding variables, symbols like (abc) mean cyclic permutations and indicate equations which are omitted.

We shall assume that at the initial instant of time the principal axis $z$ of the ellipsoid of inertia makes an angle $\theta_{0}\left(\theta_{0} \neq 1 / 2 k \pi, k=\right.$ $0,1,2$ ) with the vertical, and that the body spins about this principal axis with a high angular velocity $r_{0}$. Without loss of generality, we select the moving coordinate system in which the positive branches of the $z$-axis and of the $x$-axis do not make an obtuse angle with the direction of gravitational acceleration at the initial instant of time.

In this coordinate system, then $r_{0} \geqslant 0$, and because of the restrictions on $\theta_{0}$, the initial values $\gamma_{0}$ and $\gamma_{0}{ }^{\prime \prime}$ must satisfy the conditions

$$
\begin{equation*}
\Upsilon_{0} \geqslant 0, \quad 0<\Upsilon_{0}^{\prime \prime}<1 \tag{1.3}
\end{equation*}
$$

Let us introduce

$$
\begin{array}{ll}
a=\frac{A}{C}, & b=\frac{B}{C}, \quad c^{2}=\frac{M g l}{C}, \quad \mu=\frac{c \sqrt{\gamma_{0}{ }^{\prime \prime}}}{r_{0}}  \tag{1.4}\\
x_{0}=l x_{0}{ }^{\prime}, \quad y_{0}=l y_{0^{\prime}}, \quad z_{0}=l z_{0}{ }^{\prime}, \quad l^{2}=x_{0}{ }^{2}+y_{0}{ }^{2}+z_{0}{ }^{2}
\end{array}
$$

By assuming that $r_{0}$ is large, we assume that $\mu$ is small. Let us now introduce a new variable through the formulas $p_{1}, q_{1}, r_{1}, \gamma_{1}, \gamma_{1}{ }^{\prime}, \gamma_{1}{ }^{\prime \prime}$ and $T$

$$
\begin{align*}
& p=c \sqrt{\gamma_{0}^{\prime \prime}} p_{1}, \quad q=c \sqrt{\gamma_{0}^{\prime \prime}} q_{1}, \quad r=r_{0} r_{1}  \tag{1.5}\\
& \gamma=\gamma_{0}{ }^{\prime \prime} \gamma_{1}, \quad \gamma^{\prime}=\gamma_{0}{ }^{\prime \prime} \gamma_{1}, \quad \gamma^{\prime \prime}=\gamma_{0}{ }^{\prime \prime} \tau_{1}{ }^{\prime \prime}, \quad t=\tau / r_{0}
\end{align*}
$$

Equations (1.1) and their integrals (1.2) will, with the new variable, assume the form

$$
\begin{align*}
& \dot{p}_{1}+A_{1} q_{1} r_{1}=\mu a^{-1}\left(y_{0}{ }^{\prime} \gamma^{\prime \prime}{ }_{1}-z_{0}{ }^{\prime} \gamma_{1}{ }^{\prime}\right), \quad \gamma_{1}=r_{1} \gamma_{1}{ }^{\prime}-\mu q_{1} \gamma_{1}{ }^{\prime \prime} \\
& \dot{q}_{1}+B_{1} p_{1} r_{1}=\mu b^{-1}\left(z_{0}{ }^{\prime} \gamma_{1}-x_{0}{ }^{\prime} \gamma_{1}{ }^{\prime \prime}\right), \quad \dot{\gamma}_{1}{ }^{\prime}=\mu p_{1} \gamma_{1}{ }^{\prime \prime}-r_{1} \gamma_{1}  \tag{1.6}\\
& \dot{r}_{1}=\mu^{2}\left(-C_{1} p_{1} q_{1}+x_{0}{ }^{\prime} \Upsilon_{1}^{\prime}-y_{0}^{\prime} \Upsilon_{1}\right), \quad \dot{\gamma}_{1}^{\prime \prime}=\mu\left(q_{1} \Upsilon_{1}-p_{1} \Upsilon_{1}^{\prime}\right) \quad(\dot{u}=d u / d \tau) \\
& r_{1}{ }^{2}=1+\mu^{2} S_{1} \quad\left\{S_{1}=a\left(p_{10}{ }^{2}-p_{1}{ }^{2}\right)+b\left(q_{10}{ }^{2}-q_{1}{ }^{2}\right)-\right. \\
& -2\left[x_{0}{ }^{\prime}\left(\gamma_{10}-\gamma_{1}\right)+y_{0}{ }^{\prime}\left(\gamma_{10}{ }^{\prime}-\gamma_{1}{ }^{\prime}\right)+z_{0}{ }^{\prime}\left(1-\gamma_{1}{ }^{\prime \prime}\right)\right\}  \tag{1.7}\\
& r_{1} \gamma_{1}{ }^{\prime \prime}=1+\mu S_{2} \quad\left\{S_{2}=a\left(p_{10} \gamma_{10}-p_{1} \gamma_{1}\right)+b\left(q_{10} \Upsilon_{10}{ }^{\prime}-q_{1} \gamma_{1}\right)\right\}  \tag{1.8}\\
& \gamma_{1}{ }^{2}+\gamma_{1}{ }^{\prime 2}+\gamma_{1}{ }^{\prime 2}=\left(\gamma_{0}\right)^{{ }^{-2}} \tag{1.9}
\end{align*}
$$

Using the first integrals (1.7) and (1.8), we express the variables $r_{1}$ and $\gamma_{1}$ " in terms of the remaining variables $p_{1}, q_{1}, \gamma_{1}{ }^{\prime}, \gamma_{1} "$, and in those of their initial values $p_{10}, q_{10}, \gamma_{10}, \gamma_{10}$, and in those of the small parameter $\mu$

$$
\begin{array}{r}
r_{1}=1+1 / 2 \mu^{2}\left[S_{1}+2 z_{0}^{\prime}\left(1-\gamma_{1}^{\prime \prime}\right)\right]+\ldots \\
\gamma_{1}^{\prime \prime}=1+\mu S_{2}-1 / 2 \mu^{2}\left[S_{1}+2 z_{0}^{\prime}\left(1-\gamma_{1}^{\prime \prime}\right)\right]+\ldots \tag{1.10}
\end{array}
$$

and reduce the remaining four equations of the system (1.6) to two second order equations

$$
\begin{gather*}
\ddot{p}_{1}+\omega^{2} p_{1}=\mu\left[z_{0}^{\prime}\left(a^{-1}+A_{1} b^{-1}\right) \gamma_{1}+A_{1} b^{-1} x_{0}{ }^{\prime}\right]+\mu^{2}\left\{-\omega^{2} p_{1}\left[S_{1}+\right.\right. \\
\left.+2 z_{0}^{\prime}\left(1-\gamma_{1}^{\prime \prime}\right)\right]+A_{1} b^{-1} x_{0}^{\prime} S_{2}+A_{1} C_{1} p_{1} q_{1}^{2}-A_{1} x_{0}^{\prime} q_{1} \gamma_{1}^{\prime}-y_{0}^{\prime} a^{-1} p_{1} \gamma_{1}^{\prime}+ \\
\left.\quad+y_{0}^{\prime}\left(A_{1}+a^{-1}\right) q_{1} \gamma_{1}-z_{0}^{\prime} a^{-1} p_{1}\right\}+\mu^{3} z_{0}^{\prime}\left\{{ } ^ { 1 } I _ { 2 } ( a ^ { - 1 } - b ^ { - 1 } A _ { 1 } ) \left[S_{1}+\right.\right. \\
\left.\left.+2 z_{0}^{\prime}\left(1-\gamma_{1}^{\prime \prime}\right)\right] \gamma_{1}-\left(2 \omega^{2}+a^{-1}\right) p_{1} S_{2}\right\}+\ldots  \tag{1.11}\\
\ddot{\gamma_{1}}+\gamma_{1}=\mu\left(1+B_{1}\right) p_{1}+\mu^{2}\left\{-\gamma_{1}\left[S_{1}+2 z_{0}^{\prime}\left(1-\gamma_{1}^{\prime \prime}\right)\right]+\left(1+B_{1}\right) p_{1} S_{2}+\right. \\
\left.+\left(1+C_{1}\right) p_{1} q_{1} \gamma_{1}^{\prime}+x_{0}^{\prime} \gamma_{1}^{\prime 2}-y_{0}^{\prime} \gamma_{1} \gamma_{1}^{\prime}-b^{-1} z_{0}^{\prime} \gamma_{1}+b^{-1} x_{0}^{\prime}-q_{1}^{2} \gamma_{1}\right\}+ \\
 \tag{1.12}\\
+\mu^{3}\left[-z_{0}^{\prime}\left(2+b^{-1}\right) \gamma_{1} S_{2}+2 b^{-1} x_{0}^{\prime} S_{2}\right]+\cdots  \tag{1.13}\\
\omega^{2}=-A_{1} B_{1}=(a-1)(b-1) / a b=(A-C)(B-C) / A B
\end{gather*}
$$

Solving the first and fourth equation of system (1.6) for $q_{1}$ and $\gamma_{1}{ }^{\prime}$ we obtain
$q_{1}=\left(A_{1} r_{1}\right)^{-1}\left[-\dot{\rho}_{1}+\mu a^{-1}\left(y_{0}{ }^{\prime} \gamma_{1}{ }^{\prime \prime}-z_{0}{ }^{\prime} \gamma_{1}{ }^{\prime}\right)\right], \quad \gamma_{1}^{\prime}=\left(r_{1}\right)^{-1}\left[\dot{\gamma}_{1}+\mu q_{1} \gamma_{1}{ }^{\prime \prime}\right]$
in which $r_{1}$ and $\gamma_{1}{ }^{\prime \prime}$ are replaced by (1.10). Substituting (1.14) in the right-hand terms of equations (1.11) and (1.12), we obtain a quasilinear autonomous system with two degrees of freedom, whose right-hand sides depend on $p_{1}, \dot{p}_{1}, \gamma_{1}, \dot{\gamma}_{1}, p_{10}, \dot{p}_{10}, \gamma_{10}$ and $\dot{\gamma}_{10}$.

We want to find periodic solutions of this system under conditions $A>B>C$, or $A<B<C$ ( $\omega^{2}$ is positive). In the first case, $\omega<1$ and the $z$-axis should coincide with the major axis of the ellipsoid of inertia. In the second case, $\omega \geqslant 1$ and the $z$-axis should coincide with the minor axis of the ellipsoid of inertia. The case $\omega=1$ corresponds to the disc

$$
\begin{equation*}
A+B=C \quad(a+b=1), \quad z_{0}=0 \tag{1.15}
\end{equation*}
$$

Let us introduce new variables $p_{2}$ and $\gamma_{2}$ through

$$
\begin{equation*}
p_{1}=p_{2}+\mu \kappa+\mu \kappa_{1} \Upsilon_{2}, \quad \Upsilon_{1}=\Upsilon_{2}+\mu v p_{2} \tag{1.16}
\end{equation*}
$$

$$
\begin{equation*}
x=\frac{x_{0}{ }^{\prime} A_{1}}{b \omega)^{2}}, \quad x_{1}=-\frac{z_{0}^{\prime}}{1-\omega^{2}}\left(\frac{1}{a}+\frac{A_{1}}{b}\right), \quad v=\frac{1+B_{1}}{1-\omega^{2}} \tag{1.17}
\end{equation*}
$$

Using formulas (1.16) and (1.10) and (1.14) we can obtain the following power series expansion in $\mu$ :

$$
\begin{gather*}
S_{i}=S_{i 1}+2^{2-i} \mu S_{i 2}+\ldots \quad(i=1,2) \\
r_{1}=1+{ }^{1} / 2 \mu^{2} S_{11}+\ldots, \gamma_{1}^{\prime \prime}=1+\mu S_{21}+\mu^{2}\left(S_{22}-{ }^{1 / 2} S_{11}\right)+\ldots  \tag{1.18}\\
q_{1}=-A^{-1} \dot{p}_{2}+\mu A_{1}^{-1}\left(y_{0}^{\prime} a^{-1}-x_{2} \dot{\gamma}_{2}\right)+\ldots, \quad \gamma_{1}^{\prime}=\dot{\gamma}_{2}+\mu v_{2} \dot{p}_{2}+\ldots \\
x_{2}=x_{1}+a^{-1} z_{0}, \quad v_{2}=v-A_{1}^{-1} \tag{1.19}
\end{gather*}
$$

where

$$
\begin{gather*}
S_{11}=a\left(p_{20}{ }^{2}-p_{2}{ }^{2}\right)+b\left(\dot{p}_{20}{ }^{2}-\dot{p}_{2}{ }^{2}\right) / A_{1}{ }^{2}-  \tag{1.20}\\
-2\left[x_{0}{ }^{\prime}\left(\gamma_{20}-\gamma_{2}\right)+y_{0}{ }^{\prime}\left(\dot{\gamma}_{20}-\tau_{2}\right)\right] \\
S_{12}=a\left[x\left(p_{20}-p_{2}\right)+x_{1}\left(p_{20} \gamma_{20}-p_{2} \gamma_{2}\right)\right]-b A_{1}{ }^{-2}\left[y_{0}{ }^{\prime} a^{-1}\left(\dot{p}_{20}-\dot{p}_{2}\right)-\right. \\
\left.\quad-x_{2}\left(\dot{\gamma}_{20} \dot{p}_{20}-\dot{\gamma}_{2} \dot{p}_{2}\right)\right]-x_{0}{ }^{\prime} v\left(p_{20}-p_{2}\right)-y_{0}{ }^{\prime} v_{2}\left(\dot{p}_{20}-\dot{p}_{2}\right)+z_{0}{ }^{\prime} S_{21} \\
S_{21}=a\left(p_{20} \gamma_{20}-p_{2} \gamma_{2}\right)-b \dot{A}_{1}{ }^{-1}\left(\dot{p}_{20} \gamma_{20}-\dot{p}_{2} \dot{\gamma_{2}}\right) \\
S_{22}=a\left[v\left(p_{20}{ }^{2}-p_{2}{ }^{2}\right)+x\left(\gamma_{20}-\gamma_{2}\right)+x_{1}\left(\gamma_{20}{ }^{2}-\gamma_{2}{ }^{2}\right)\right]+ \\
\quad+b A_{1}{ }^{-1}\left[-v_{2}\left(\dot{p}_{20}{ }^{2}-\dot{p}_{2}{ }^{2}\right)+a^{-1} y_{0}{ }^{\prime}\left(\dot{\gamma}_{20}-\dot{\gamma}_{2}\right)-x_{2}\left(\dot{\gamma}_{20}{ }^{2}-\dot{\gamma}_{2}{ }^{2}\right)\right]
\end{gather*}
$$

Substituting (1.16) and (1.18) into equations (1.11) and (1.12), we obtain

$$
\begin{gather*}
\ddot{p}_{2}+\omega^{2} p_{2}=\mu^{2} F\left(p_{2}, \dot{p}_{2}, \gamma_{2}, \dot{\gamma}_{2}, \mu\right), \quad F=F_{2}+\mu F_{3}+\ldots  \tag{1.21}\\
\ddot{\gamma}_{2}+\gamma_{2}=\mu^{2} \Phi\left(p_{2}, \dot{p}_{2}, \gamma_{2}, \dot{\gamma}_{2}, \mu\right) \quad \Phi=\Phi_{2}+\mu \Phi_{3}+\ldots \\
F_{2}=f_{2}-v x_{1}\left(1-\omega^{2}\right) p_{2} \\
F_{3}=f_{3}-x_{1} \varphi_{2}-v x_{1} \nless\left(1-\omega^{2}\right)-v x_{1}^{2}\left(1-\omega^{2}\right) \gamma_{2} \tag{1.22}
\end{gather*}
$$

$\Phi_{2}=\varphi_{2}+\nu x\left(1-\omega^{2}\right)+\nu x_{1}\left(1-\omega^{2}\right) \gamma_{2}, \quad \Phi_{3}=\varphi_{3}-v f_{2}+\nu^{2} \varkappa_{1}\left(1-\omega^{2}\right) p_{2}$ where

$$
\begin{gather*}
f_{2}=-\omega^{2} p_{2} S_{11}+A_{1} b^{-1} x_{0}{ }^{\prime} S_{21}+C_{1} A_{1}{ }^{-1} p_{2} \dot{p}_{2}{ }^{2}+x_{0}{ }^{\prime} \dot{p}_{2} \dot{\gamma}_{2}-y_{0}{ }^{\prime} a^{-1} p_{2} \dot{\gamma}_{2}-  \tag{1.23}\\
\quad-y_{0}{ }^{\prime} A_{1}{ }^{-1}\left(A_{1}+a^{-1}\right) \Upsilon_{2} \dot{p}_{2}-z_{0}^{\prime} a^{-1} p_{2} \\
f_{3}=-\omega^{2}\left(x S_{11}+x_{1} \gamma_{2} S_{11}+2 p_{2} S_{12}\right)+A_{1} b^{-1} x_{0}{ }^{\prime} S_{22}+C_{1} A_{1}{ }^{-1}\left[x \dot{p}_{2}{ }^{2}+\right. \\
\left.+x_{1} \gamma_{2} \dot{p}_{2}{ }^{2}-2 p_{2} \dot{p}_{2}\left(y_{0}{ }^{\prime} a^{-1}-x_{2} \dot{\gamma}_{2}\right)\right]-x_{0}{ }^{\prime}\left[-v_{2} \dot{p}_{2}{ }^{2}+\dot{\gamma}_{2}\left(y_{0}{ }^{\prime} a^{-1}-x_{2} \dot{\gamma}_{2}\right)\right]- \\
-y_{0}{ }^{\prime} a^{-1}\left[\dot{\gamma}_{2}\left(x+x_{1} \gamma_{2}\right)+v_{2} p_{2} \dot{p}_{2}\right]+y_{0}^{\prime} A_{1}{ }^{-1}\left(A_{1}+a^{-1}\right)\left[\gamma_{2}\left(y_{0}{ }^{\prime} a^{-1}-x_{2}\right)-\right. \\
\left.-v p_{2} \dot{p}_{2}\right]-z_{0}{ }^{\prime} a^{-1}\left(x+x_{1} \gamma_{2}\right)+{ }_{2} 2_{0} z_{0}^{\prime}\left(a^{-1}+A_{1} b^{-1}\right) \gamma_{2} S_{11}+z_{0}{ }^{\prime} a^{-1} p_{2} S_{21}
\end{gather*}
$$

$$
\begin{gathered}
\varphi_{2}=-\gamma_{2} S_{11}+\left(1+B_{1}\right) p_{2} S_{21}-\left(1-C_{1}\right) A_{1}{ }^{-1} p_{2} \dot{p}_{2} \dot{\gamma}_{2}+x_{0} \dot{\gamma}_{2}{ }^{2}-y_{0}{ }^{\prime} \gamma_{2} \dot{\gamma}_{2}- \\
-z_{0}^{\prime} b^{-1} \gamma_{2}+x_{0}{ }^{\prime} b^{-1}-A_{1}{ }^{-2} \gamma_{2} \dot{p}_{2}{ }^{2} \\
\varphi_{3}=-v p_{2} S_{11}-2 \gamma_{2} S_{21}+\left(1+B_{1}\right) p_{2} S_{22}+\left(1+B_{1}\right)\left(x+x_{1} \gamma_{2}\right) S_{21}+ \\
+\left(1-C_{1}\right) A_{1}{ }^{-1}\left[-\left(x+x_{1} \gamma_{2}\right) \dot{p}_{2} \dot{\gamma}_{2}-v_{2} p_{2} \dot{p}_{2}{ }^{2}+p_{2} \dot{\gamma}_{2}\left(y_{0}{ }^{\prime} a^{-1}-x_{2} \dot{\gamma}_{2}\right)\right]+ \\
+2 x_{0}{ }^{\prime} v_{2} \dot{p}_{2} \dot{\gamma}_{2}-y_{0}^{\prime}\left(v p_{2} \gamma_{2}+\gamma_{2} \gamma_{2} \dot{p}_{2}\right)-z_{0}^{\prime} b^{-1} v p_{2}-z_{0} b^{-1} \gamma_{2} S_{21}+ \\
+2 x_{0} b^{-1} S_{21}-A_{1}{ }^{-2}\left[-2 \gamma_{2} \dot{p}_{2}\left(y_{0}^{\prime} a^{-1}-x_{2} \dot{\gamma}_{2}\right)+v p_{2} \dot{p}_{2}{ }^{2}\right]
\end{gathered}
$$

This system has a first integral which can be obtained from the integral (1.9)

$$
\begin{equation*}
\gamma_{2}^{2}+\dot{\tau}_{12}^{2}+2 \mu\left(v p_{2} \gamma_{2}+v_{2} \dot{p}_{2} \tilde{\tau}_{2}+S_{21}\right)+\mu^{2}(\ldots)=\left(\gamma_{0}{ }^{\prime \prime}\right)^{-2}-1 \tag{1.24}
\end{equation*}
$$

We are going to look for periodic solutions $p_{2}(T, \mu), \dot{p}_{2}(T, \mu)$, $\gamma_{2}(\tau, \mu)$ and $\dot{\gamma}_{2}(\tau, \mu)$ of system (1.21), but only those which satisfy conditions

$$
\begin{equation*}
p_{2}(0,0)=0, \quad \dot{p}_{2}(0,0)=0, \quad \dot{\gamma}_{2}(0, \mu)=0 \tag{1.25}
\end{equation*}
$$

Since system (1.21) is autonomous, the above conditions do not affect the generality of the solution.

Since the frequencies of the generating system

$$
\begin{equation*}
\ddot{p}_{2}+\omega^{2} p_{2}=0, \quad \ddot{\gamma}_{2}+\tau_{2}=0 \tag{1.26}
\end{equation*}
$$

are $\omega$ and 1 , we can construct [5] periodic solutions of system (1.21) in three different ways:

1) When the two frequencies are distinct but commensurate ( $\omega=m / n$, where $m$ and $n$ are relative primes);
2) When the two frequencies are equal ( $\omega=1$ );
3) When the two frequencies are noncomensurate ( $\omega$ is irrational),
2. Let us consider the first case, $\omega=m / n$. In this case, there exist periodic solutions of the generating system (1.26) with the period $T_{0}=2 \pi n$

$$
\begin{equation*}
p_{2}^{(0)}=M_{1} \cos \omega \tau+M_{2} \sin \omega \tau, \quad \gamma_{2}^{(0)}=M_{3} \cos \tau \tag{2.1}
\end{equation*}
$$

We assume that the initial autononous system (1.21) has periodic solutions with the period $T_{0}+\alpha$, which reduce to the generating solution (2.1) when $\mu=0$. We shall write the initial conditions through
the relations

$$
\begin{gather*}
p_{2}(0, \mu)=M_{1}+\beta_{1}, \quad \dot{p}_{2}(0, \mu)=\omega\left(M_{2}+\beta_{2}\right) \\
\gamma_{2}(0, \mu)=M_{3}+\dot{\beta}_{3}, \quad \dot{\gamma}_{2}(0, \mu)=0 \tag{2.2}
\end{gather*}
$$

Let us also introduce the operator

$$
U=u+\frac{\partial u}{\partial M_{1}} \beta_{1}+\frac{\partial u}{\partial M_{2}} \beta_{2}+\frac{\partial u}{\partial M_{3}} \beta_{3}+\frac{1}{2} \frac{\partial^{2} u}{\partial M_{1}{ }^{2}} \beta_{1}{ }^{2}+\ldots \quad\binom{U=G_{k}, H_{k}}{u=g_{k}, h_{k}}
$$

and express [5] the periodic solution which we want to find in the form

$$
\begin{gather*}
p_{2}(\tau, \mu)=\left(M_{1}+\beta_{1}\right) \cos \omega \tau+\left(M_{2}+\beta_{2}\right) \sin \omega \tau+\sum_{k=2}^{\infty} G_{k}(\tau) \mu^{k} \\
\gamma_{2}(\tau, \mu)=\left(M_{3}+\beta_{3}\right) \cos \tau+\sum_{k=2}^{\infty} H_{k}(\tau) \mu^{k}  \tag{2.3}\\
g_{k}(\tau)=\frac{1}{\omega} \int_{0}^{\tau} F_{k}^{\prime}\left(t_{1}\right) \sin \omega\left(\tau-t_{1}\right) d t_{1}, \quad h_{k}(\tau)=\int_{0}^{\tau} \Phi_{k}^{\prime}\left(t_{1}\right) \sin \left(\tau-t_{1}\right) d t_{1} \\
F_{k}^{\prime}(\tau)=\frac{1}{(k-2)!}\left(\frac{d^{k-2} F}{d \mu^{k-2}}\right)_{\beta=\mu=0}, \quad \Phi_{k}^{\prime}(\tau)=\frac{1}{(k-2)!}\left(\frac{d^{k-2} \Phi}{d \mu^{k-2}}\right)_{\beta=\mu=0} \quad(2.4) \tag{2.4}
\end{gather*}
$$

Here $\beta_{1}, \omega \beta_{2}$ and $\beta_{3}$ are deviations of the initial values of $p_{2}, \dot{p}_{2}$ and $\gamma_{2}$ in the periodic solutions of equations (1.21) from the initial values of the same quantities in the generating solution (2.1). These deviations are functions of $\mu$, and they vanish when $\mu=0$.

Since the right-hand sides in system (1.21) begin from a term of the order $\mu^{2}$, we have the following relations:

$$
\begin{align*}
& F_{k}^{\prime}(\tau)=F_{k}\left(p_{2}^{(0)}, \dot{p}_{2}^{(0)}, \gamma_{2}^{(0)}, \dot{\gamma}_{2}^{(0)}\right) \equiv F_{k}^{(0)} \\
& \Phi_{k}^{\prime}(\tau)=\Phi_{k}\left(p_{2}^{(0)}, \dot{p}_{2}^{(0)}, \gamma_{2}^{(0)}, \dot{\gamma}_{2}^{(0)}\right) \equiv \Phi_{k}^{(0)} \tag{2.5}
\end{align*} \quad(k=2,3)
$$

Let us find now the functions $F_{2}^{(0)}$ and $\Phi_{2}^{(0)}$. Introducing the notation

$$
\begin{equation*}
E=\sqrt{M_{1}^{2}+M_{2}^{2}}, \quad \cos \varepsilon=M_{1} / E, \quad \sin \varepsilon=M_{2} / E \tag{2.6}
\end{equation*}
$$

formulas (2.1) take the form

$$
\begin{equation*}
p_{2}^{(0)}(\tau)=E \cos (\omega \tau-\varepsilon), \quad \tau_{2}^{(0)}(\tau)=M_{3} \cos \tau \tag{2.7}
\end{equation*}
$$

Using (2.7), and by (1.20), formulas

$$
S_{11}^{(0)}=S_{11}\left(p_{2}^{(0)}, \quad \dot{p}_{2}^{(0)}, \gamma_{2}^{(0)}, \dot{\gamma}_{2}^{(0)}\right), \quad S_{21}^{(0)}=S_{21}\left(p_{2}^{(0)}, \dot{p}_{2}^{(0)}, \tau_{2}^{(0)}, \dot{\gamma}_{2}^{(0)}\right)
$$

become

$$
\begin{gather*}
S_{11}{ }^{(0)}=E^{2}\left\{\left[a\left(\cos ^{2} \varepsilon-1 / 2\right)+b_{\omega^{2} A_{1}{ }^{-2}\left(\sin ^{2} \varepsilon-1 / 2\right)+}^{\left.+1 / 2\left(b \omega^{2} A_{1}{ }^{-2}-a\right) \cos 2(\omega \tau-\varepsilon)\right\}-2 M_{3}\left[x_{0}{ }^{\prime}(1-\cos \tau)+y_{0}{ }^{\prime} \sin \tau\right]}\right.\right. \\
S_{21}{ }^{(0)}=M_{3} E\left\{a \cos \varepsilon+1_{2}\left(b \omega A_{1}^{-1}-a\right) \cos [(\omega-1) \tau-\varepsilon]-\right.  \tag{2.8}\\
\left.-12_{2}\left(b \omega A_{1}{ }^{-1}+a\right) \cos [(\omega+1) \tau-\varepsilon]\right\}
\end{gather*}
$$

Substituting (2.7) and (2.8) into formulas (1.22) and (1.23), respectively, we obtain for all the values of $\omega$ (with the exception of the case $\omega=1 / 2$, which will be considered later) the expressions
$F_{2}{ }^{(0)}=M_{1} L(\omega) \cos \omega \tau+M_{2} L(\omega) \sin \omega \tau+\ldots, \quad \Phi_{2}{ }^{(0)}=M_{3} N(\omega) \cos \tau+\ldots$ where

$$
\begin{align*}
L(\omega)= & \omega^{2}\left[-\left(a M_{1}^{2}+b \omega^{2} A_{1}{ }^{-2} M_{2}{ }^{2}\right)+1 / 4\left(M_{1}^{2}+M_{2}{ }^{2}\right)\left(C_{1} A_{1}^{-1}+\right.\right. \\
& \left.\left.+3 a+b^{2} A_{1}{ }^{-2}\right)\right]+2 M_{3} \omega^{2} x_{0}^{\prime}-\left[z_{0}^{\prime} a^{-1}+\chi_{1} v\left(1-\omega^{2}\right)\right] \\
N(\omega)= & -\left(a M_{1}{ }^{2}+b_{\left.\omega^{2} A_{1}{ }^{-2} M_{2}{ }^{2}\right)-1 / 2\left(M_{1}{ }^{2}+M_{2}^{2}\right)\left[a B_{1}+\right.}+\omega^{2} A_{1}^{-2}(1-b)\right]+2 M_{3} x_{0}^{\prime}-\left[z_{0}^{\prime} b^{-1}-x_{1} v\left(1-\omega^{2}\right)\right] \tag{2.10}
\end{align*}
$$

From formulas (2.4) and (2.5) we obtain

$$
\begin{equation*}
g_{k}\left(T_{0}\right)=-\frac{1}{\omega} \int_{0}^{T_{0}} F_{k}^{(0)}\left(t_{1}\right) \sin \omega t_{1} d t_{1}, \quad \dot{g_{k}}\left(T_{0}\right)=\int_{0}^{T_{0}} F_{k}^{(0)}\left(t_{1}\right) \cos \omega t_{1} d t_{1} \tag{2.11}
\end{equation*}
$$

$h_{k}\left(T_{0}\right)=-\int_{0}^{T_{0}} \Phi_{k}{ }^{(0)}\left(t_{1}\right) \sin t_{1} d t_{1}, \quad \dot{h}_{k}\left(T_{0}\right)=\int_{0}^{T_{0}} \Phi_{k}{ }^{(0)}\left(t_{1}\right) \cos t_{1} d t_{1}\binom{T_{0}=2 \pi n}{k=2,3}$
Hence, by using (2.9) we have

$$
\begin{gather*}
g_{2}\left(T_{0}\right)=-\pi n_{\omega}^{-1} M_{2} L(\omega), \quad \dot{g}_{2}\left(T_{0}\right)=\pi n M_{1} L(\omega)  \tag{2.12}\\
h_{2}\left(T_{0}\right)=0, \quad \dot{h_{2}}\left(T_{0}\right)=\pi n M_{3} N(\omega)
\end{gather*}
$$

The constants $H_{1}, \omega H_{2}$ and $M_{3}$, which represent the initial conditions of the generating solution (2.1), the deviations $\beta_{1}(\mu), \omega \beta_{2}(\mu)$ and $\beta_{3}(\mu)$ and the correction for the period $\alpha$, must be found from the conditions for periodicity of the solutions $p_{2}(T, \mu), \gamma_{2}(\tau, \mu)$ and their first derivatives; these conditions have the form

$$
\begin{array}{ll}
\Psi_{1}=p_{2}\left(T_{0}+\alpha, \mu\right)-p_{2}(0, \mu)=0, & \Psi_{2}=\dot{p}_{2}\left(T_{0}+\alpha, \mu\right)-\dot{p}_{2}(0, \mu)=0  \tag{2.13}\\
\Psi_{3}=\gamma_{2}\left(T_{0}+\alpha, \mu\right)-\gamma_{2}(0, \mu)=0, & \Psi_{4}=\dot{\gamma}_{2}\left(T_{0}+\alpha, \mu\right)-\dot{\gamma}_{2}(0, \mu)=0
\end{array}
$$

However, on the strength of the existence of the first integral (1.24) of system (1.21), the condition for periodicity $\Psi_{3}=0$ is not independent [6]. Indeed, writing the integral (1.24) in the form

$$
\gamma_{2}^{2}\left(T_{0}+\alpha, \mu\right)+\dot{\gamma}_{2}^{2}\left(T_{0}+\alpha, \mu\right)+\mu(\ldots)=\gamma_{2}^{2}(0, \mu)+\dot{\gamma}_{2}^{2}(0, \mu)+\mu(\ldots)
$$

and substituting formulas (2.13), we obtain, according to condition (2.2)

$$
\begin{equation*}
2\left(M_{3}+\beta_{3}\right) \Psi_{3}+\Psi_{3}^{2}+\mu \varphi_{1}\left(\Psi_{1}, \Psi_{2}, \Psi_{3}, \Psi_{4}, \mu\right)=0 \tag{2.14}
\end{equation*}
$$

in which $\varphi_{1}$ is an entire function of all its arguments, and, besides, $\varphi_{1}(0,0,0, \mu)=0$. When $M_{3} \neq 0$, which, as is shown later, is always the case, then by formula (2.14) $\Psi_{3}=f\left(\Psi_{1}, \Psi_{2}, \Psi_{4}, \mu\right)$, where $f$ is an entire function of all its arguments, and $f(0,0,0, \mu)=0$. Consequentl we can see from (2.13) that condition $\Psi_{3}=0$ is automatically satisfied when the remaining conditions

$$
\begin{equation*}
\Psi_{1}=\Psi_{2}=\Psi_{4}=0 \tag{2.15}
\end{equation*}
$$

are satisfied.
Substituting the initial conditions (2.2) in the integral (1.24) when $T=0$, we obtain the equations determining $M_{3}$ and $\rho_{3}$

$$
M_{3}^{2}+2 M_{3} \beta_{3}+\beta_{3}^{2}+2 \mu v M_{3}\left(M_{1}+\beta_{1}\right)+\ldots=\left(\gamma_{0}^{n}\right)^{-2}-1
$$

Assuming that $\gamma_{0}{ }^{\prime \prime}$ is independent of $\mu$, we obtain

$$
M_{3}^{2}=\left(\gamma_{0}^{\prime \prime}\right)^{-2}-1, \quad \beta_{3}^{2}+2 M_{3} \beta_{3}+2 \mu v M_{3}\left(M_{1}+\beta_{1}\right)+\ldots=0(2.16)
$$

From equations (2.16) and from condition (1.3) it follows that (2.17)

$$
0<M_{3}=\left(1-\gamma_{0}^{\prime \prime 2}\right)^{1 / 2}\left(\gamma_{0}^{\prime \prime}\right)^{-1}<\infty, \quad \beta_{3}=-\mu_{v}\left(M_{1}+\beta_{1}\right)+\ldots
$$

because $\gamma_{0}{ }^{\prime \prime}$ is an arbitrary parameter, and $H_{3}$ is an arbitrary positive constant.

In this way, the periodic solution (2.3) depends on one arbitrary constant $M_{3}$ and on a function $\beta_{3}(\mu)$, vanishing when $\mu=0$. This property does not depend on the form of $\alpha$ and occurs in all considered cases.

Expanding the right-hand sides of equations (2.13) in power series of $\alpha$ and retaining only the linear terms (neglecting even the terms $\mu^{2} \alpha$ ), we obtain the independent conditions of periodicity (2.15) in the form

$$
\begin{gather*}
p_{2}\left(T_{0}, \mu\right)-M_{1}-\beta_{1}+\alpha \omega\left(M_{2}+\beta_{2}\right)=0 \\
\dot{p}_{2}\left(T_{0}, \mu\right)-\omega\left(M_{2}+\beta_{2}\right)-\alpha \omega^{2}\left(M_{1}+\beta_{1}\right)=0  \tag{2.18}\\
\dot{\gamma}_{2}\left(T_{0}, \mu\right)-\alpha\left(M_{3}+\beta_{3}\right)=0
\end{gather*}
$$

By the last equation, from (2.18) and by (2.17) and (2.3), we determine the function

$$
\begin{equation*}
\alpha=\mu^{2}\left[\dot{H}_{2}\left(T_{0}\right)+\mu \dot{H}_{3}\left(T_{0}\right)+\ldots\right] /\left(M_{3}+\beta_{3}\right) \quad\left(T_{0}=2 \pi n\right) \tag{2.19}
\end{equation*}
$$

We can conclude from it that by neglecting terms of the order $\alpha^{2}$ and $\mu^{2} \alpha$ in system (2.18), we neglect also terms of the order $\mu^{4}$.

Using (1.25) and (2.2) we shall now investigate those periodic solutions which arise when the fundamental amplitudes vanish, that is

$$
\begin{equation*}
M_{1}=0, \quad M_{2}=0 \tag{2.20}
\end{equation*}
$$

Substituting relations (2.19), (2.20) and (2.3) in the first two equations (2.18) and cancelling $\mu^{2}$, we obtain the system determining $\rho_{1}$ and $\beta_{2}$

$$
\begin{aligned}
G_{2}\left(T_{0}\right)+\mu G_{3}\left(T_{0}\right)+\omega \beta_{2} & {\left[\dot{H}_{2}\left(T_{0}\right)+\mu \dot{H}_{3}\left(T_{0}\right)+\ldots\right] /\left(M_{3}+\beta_{3}\right)+} \\
& +\mu^{2}(\ldots)=0 \\
\dot{G}_{2}\left(T_{0}\right)+\mu \dot{G}_{3}\left(T_{0}\right)-\omega^{2} \beta_{1} & {\left[\dot{H}_{2}\left(T_{0}\right)+\mu \dot{H}_{3}\left(T_{0}\right)+\ldots\right] /\left(M_{3}+\beta_{3}\right)+} \\
& +\mu^{2}(\ldots)=0
\end{aligned}
$$

Using (2.12) we can transform the above system into

$$
\begin{gather*}
-\left[L_{1}(\omega)-\omega^{2} N_{1}(\omega)\right] \pi n \beta_{2} / \omega+\mu\left[G_{3}\left(T_{0}\right)+\ldots\right]=0  \tag{2.29}\\
{\left[L_{1}(\omega)-\omega^{2} N_{1}(\omega)\right] \operatorname{sn} \beta_{1}+\mu\left[\dot{G}_{3}\left(T_{0}\right)+\ldots\right]=0}
\end{gather*}
$$

Here $L_{1}(\omega)$ and $N_{1}(\omega)$ are obtained from (2.10) by replacing $M_{1}, M_{2}$ and $M_{3}$ by $\beta_{1}, \beta_{2}$ and $M_{3}+\beta_{3}$. By (1.13), (1.17) and (1.19), we obtain

$$
\begin{gather*}
L_{1}(\omega)-\omega^{2} N_{1}(\omega)=W(\omega)\left(\beta_{1}^{2}+\beta_{2}^{2}\right)-z_{0}^{\prime} W_{1}(\omega)  \tag{2.23}\\
W(\omega)=(a-1)(a+b-2) / 2 b, \quad W_{1}(\omega)=(3 a+3 b-4 a b-2) / a b \tag{2.24}
\end{gather*}
$$

From the condition that the $z$-axis has to be directed along the major or the minor axis of the ellipsoid of inertia of the body, it follows that for all the considered $\omega$, "/ $(\omega)>0$. We shall now show that
for every value of $\omega \geqslant 1$ there exists only one pair of numbers $a^{*}$ and $b^{*}$ satisfying the condition $W_{1}(\omega)=0$, and that for the values of $\omega<1$, $W_{1}(\omega) \neq 0$.

Eliminating $a^{*}$ from the system of two nonlinear equations

$$
\begin{equation*}
3 a^{*}+3 b^{*}-4 a^{*} b^{*}-2=0, \quad \omega^{2}-\left(a^{*}-1\right)\left(b^{*}-1\right) / a^{*} b^{*}=0 \tag{2.25}
\end{equation*}
$$

we find

$$
\begin{equation*}
\omega^{2}=\frac{\left(1-b^{*}\right)^{2}}{3 b^{*}\left(2 / 3-b^{*}\right)}, \quad \text { or } \quad b_{1,2}^{*}=\frac{1+\omega^{2} \pm \sqrt{\omega^{2}}\left(\omega^{2}-1\right)}{3 \omega^{2}+1} \tag{2.26}
\end{equation*}
$$

It is seen that the above formula is valid only when $\omega \geqslant 1$. Consequently, we have the following relation:
$0<a^{*} \leqslant b^{*}<1 \quad\left(a^{*}=\frac{1+\omega^{2}-\sqrt{\omega^{2}\left(\omega^{2}-1\right)}}{3 \omega^{2}+1}, \quad b^{*}=\frac{1+\omega^{2}+\sqrt{\omega^{2}\left(\omega^{2}-1\right)}}{3 \omega^{2}+1}\right)$
When $\omega=1$, it follows from formulas (2.27) that $a^{*}=b^{*}=1 / 2$, and for every value of $\omega>1$ there exists a unique pair of distinct values of $a^{*}$ and $b^{*}$ satisfying (2.25). Besides, by (2.26)

$$
\begin{equation*}
1 / 2<a^{*}<b^{*}<2 / 3 \tag{2.28}
\end{equation*}
$$

Assuming that

$$
\begin{equation*}
z_{0}^{\prime} W_{1}(\omega) \neq 0 \tag{2.29}
\end{equation*}
$$

we obtain from equations (2.22) the expansions of $\beta_{1}$ and $\beta_{2}$ in power series in $\mu$.

To make an estimate of the order of the first terms of these expansions, we shall consider the quantities $C_{3}\left(T_{0}\right)$ and $\dot{C}_{3}\left(T_{0}\right)$ under conditions (2.20). Calculations show that $C_{3}\left(T_{0}\right)=\dot{C}_{3}\left(T_{0}\right)=0$ when $\omega \neq 2$. This means [7] that the expansions of $\beta_{1}$ and $\beta_{2}$ in series of integral powers of $\mu$ begin from terms whose order is not higher than $\mu^{2}$. Consequently, when $\omega$ is rational and does not equal $1,2,1 / 2$, the first terms in the expansion of the periodic solution of the system (1.6) and the quantities $\alpha$ both under the conditions $z_{0}^{\prime} \neq 0, a=a^{*}, b=b^{*}$, can be expressed in the following form:

$$
\begin{gather*}
p_{1}=-\mu x_{0}^{\prime} / b B_{1}+\mu x_{1} M_{3} \cos \tau+\ldots, q_{1}=\mu y_{0}{ }^{\prime} / a A_{1}+\mu x_{2} A_{1}{ }^{-1} M_{3} \sin \tau+\ldots \\
r_{1}=1-\mu^{2} M_{3}\left[x_{0}^{\prime}(1-\cos \tau)+y_{0}^{\prime} \sin \tau\right]+\ldots, \gamma_{1}=M_{3} \cos \tau+\ldots \\
\gamma_{1}^{\prime}=-M_{3} \sin \tau+\ldots  \tag{2.30}\\
\tau_{1}^{\prime \prime}=1+\mu^{2}\left\{\left[M_{3}(1-a)^{-1} x_{0}^{\prime}+M_{3}^{2} z_{0}^{\prime}\left(A_{1}-1\right)\left(A_{1}+1\right)^{-1}\right]+\right. \\
\quad+M_{3}(1-b)^{-1} y_{0}^{\prime} \sin \tau-M_{3}(1-a)^{-1} x_{0}^{\prime} \cos \tau-
\end{gather*}
$$

$$
\begin{align*}
& \left.-{ }^{1} / 2 M_{3}^{2} z_{0}^{\prime}\left(A_{1}-1\right)\left(A_{1}+1\right)^{-1} \cos 2 \tau\right\}+\ldots \\
& \alpha=2 \mu^{2} \pi n\left(M_{3} x_{0}^{\prime}-z_{0}^{\prime}+\ldots\right) \tag{2.31}
\end{align*}
$$

3. Let us consider periodic solutions of system (1.21) when $\omega=2$. For this purpose we shall find $g_{3}\left(T_{0}\right)$ and $\dot{g}_{3}\left(T_{0}\right)$ under condition (2, 20). Under the last conditions formula (1.20) will be rewritten as

$$
\begin{gathered}
S_{11}^{(0)}=-2 M_{3}\left[x_{0}^{\prime}(1-\cos \tau)+y_{0}{ }^{\prime} \sin \tau\right], \quad S_{12}^{(0)}=S_{21}^{(0)}=0 \\
S_{22}{ }^{(0)}=\left[a x M_{3}+1 / 2 M_{3}{ }^{2}\left(a x_{1}+b A_{1}{ }^{-1} x_{2}\right)\right]+a^{-1} b A_{1}{ }^{-1} y_{0} M_{3} \sin \tau- \\
\quad-a x M_{3} \cos \tau-1 / 2 M_{3}^{2}\left(a x_{1}+b A_{1}^{-1} x_{2}\right) \cos 2 \tau
\end{gathered}
$$

Substituting expression (3.1) into formulas (1.23) and (1.22), and retaining only terms with $\sin 2 T$ and $\cos 2 T$, we obtain the relation

$$
\begin{equation*}
F_{3}^{(0)}=V_{1} \cos -2 \tau+V_{2} \sin 2 \tau+\ldots \tag{3.2}
\end{equation*}
$$

Here

$$
\begin{array}{r}
V_{1}=-\frac{x_{0}{ }^{\prime} z_{0}{ }^{\prime} M_{3}{ }^{2}}{6 a b^{2}}\left(12 b^{2}-b-1\right), \quad V_{2}=-\frac{y_{0}^{\prime} z_{0}{ }^{\prime} M_{3}{ }^{2}}{6 a^{2} b(1-b)}\left(9 \mathrm{a} b^{2}-17 a b+\right. \\
\left.+2 b^{2}+4 a-3 b+1\right) \tag{3.3}
\end{array}
$$

By (3.2) and (2,11) we obtain

$$
\begin{equation*}
g_{3}(2 \pi)=-1 / 2 \pi V_{2}, \quad g_{3}(2 \pi)=\pi V_{1} \quad\left(T_{0}=2 \pi\right) \tag{3.4}
\end{equation*}
$$

Let us now consider formulas (3.3). Substituting in (3.3) $a=(1-b) /$ $(3 b+1)$ obtained from (1.13) when $\omega=2$, we have

$$
\begin{align*}
& V_{1}=-6 M_{3}^{2} x_{0}^{\prime} z_{0}^{\prime} b^{-2}(1-b)^{-1}(b-1 / 3)(b+1 / 4)(b+1 / 3)  \tag{3.5}\\
& V_{2}=-3 / 2 M_{3}{ }^{2} y_{0}{ }^{\prime} z_{0}^{\prime} b^{-1}(1-b)^{-2}(b-1 / 3)(b+5)(b+1 / 3)
\end{align*}
$$

From formula (3.5) and from (1.13) it follows that $V_{1}$ and $V_{2}$ vanish simultaneously if one of the following two conditions:

$$
\begin{equation*}
a=b={ }^{1} / 3, \quad x_{0}^{\prime}=y_{0}^{\prime}=0 \tag{3.6}
\end{equation*}
$$

is satisfied (from (2.22) it follows that $z_{0}{ }^{\prime} \neq 0$ ).
This means that, if any one of the two conditions (3.6) is satisfied, then the case considered reduces to the previous one and the first terms of the expansions of the periodic solution of system (1.6) and the $\alpha$ can be obtained from formulas (2.30) and (2.31).

If $V_{1}$ and $V_{2}$ do not vanish simultaneously, then by substituting formulas (3.4) into system (2.22) and under conditions $W_{1}(2) \neq 0$ and $\omega=2(n=1)$, we obtain

$$
\beta_{1}=\mu V_{1} / z_{0}^{\prime} W_{1}(2)+\ldots, \beta_{2}=\mu V_{2} / z_{0}^{\prime} W_{1}(2)+\ldots
$$

Consequently, in formulas (2,30) only the expressions for $p_{1}, q_{1}$ and $\gamma_{1}$ "will change and become

$$
\begin{gather*}
p_{1}=-\mu \frac{x_{0}^{\prime}}{b B_{1}}+\mu \mu_{1} M_{3} \cos \tau+\mu \frac{V_{1}}{z_{0}^{\prime} W_{1}(2)} \cos 2 \tau+\mu \frac{V_{2}}{z_{0}^{\prime} W_{1}(2)} \sin 2 \tau+\ldots  \tag{3.7}\\
q_{1}=\mu \frac{y_{0}^{\prime}}{a A_{1}}+\mu \frac{x_{2} M_{3}}{A_{1}} \sin \tau+\mu \frac{2 V_{1}}{A_{1} z_{0}^{\prime} W_{1}(2)} \sin 2 \tau-\mu \frac{2 V_{2}}{A_{1} z_{0}^{\prime} W_{1}(2)} \cos 2 \tau+\ldots \\
\gamma_{1}^{\prime \prime}=1+\mu^{2} M_{5}\left(s_{0}+s_{1} \sin \tau+s_{2} \cos \tau+s_{3} \cos 2 \tau+s_{4} \cos 3 \tau+s_{5} \sin 3 \tau\right)+\ldots
\end{gather*}
$$

where

$$
\begin{gathered}
s_{0}=\frac{x_{0}^{\prime}}{1-a}+\frac{1}{2} M_{3} z_{0}^{\prime} \frac{A_{1}-1}{A_{1}+1}+\frac{a V_{1}}{z_{0}^{\prime} W_{1}(2),}, \quad s_{1}=\frac{y_{0}^{\prime}}{1-b}+\frac{V_{2}}{2 z_{0}^{\prime} W_{1}(2)}\left(\frac{2 b}{A_{1}}-a\right) \\
s_{2}=-\frac{x_{0}^{\prime}}{1-a}+\frac{V_{1}}{2 z_{0}^{\prime} W_{1}(2)}\left(\frac{2 b}{A_{1}}-a\right), \quad s_{3}=-\frac{1}{2} M_{3} z_{0}^{\prime} \frac{A_{1}-1}{A_{1}+1} \\
s_{1}=-\frac{V_{1}}{2 z_{0} W_{1}(2)}\left(\frac{2 b}{A_{1}}+a\right), \quad s_{5}=-\frac{V_{2}}{2 z_{0}^{\prime} W_{1}(2)}\left(\frac{2 b}{A_{1}}+a\right)
\end{gathered}
$$

Under conditions $W_{1}(2)=0\left(a^{*}=1 / 13\left(5-V_{12}\right)\right.$ and $b^{*}=1 / 13\left(5+V_{12}\right)$ system (2.22) will take the form

$$
\begin{equation*}
\beta_{1}\left(\beta_{1}^{2}+\beta_{2}^{2}\right)=-\mu V_{1}^{*} / W^{*}(2)+\ldots, \beta_{2}\left(\beta_{1}^{2}+\beta_{2}^{2}\right)=-\mu V_{2}^{*} / W^{*}(2)+\ldots \tag{3.8}
\end{equation*}
$$

Hence

$$
\begin{array}{ll}
\beta_{1}=\mu^{1 / 3} Y_{1}(\mu), & Y_{1}(\mu)=-V_{1}^{*}\left[W^{*}(2)\left(V_{1}^{* 2}+V_{2}^{* 2}\right)\right]^{-1 / 3}+\mu^{1 / 3}(\ldots)  \tag{3.9}\\
\beta_{2}=\mu^{1 / 3} Y_{2}(\mu), & Y_{2}(\mu)=-V_{2}^{*}\left[W^{*}(2)\left(V_{1}^{* 2}+V_{2}^{* 2}\right)\right]^{-1 / 3}+\mu^{1 / 3}(\ldots)
\end{array}
$$

In this case the first terms of the expansion of the periodic solution of system (1.6) and the quantity $\alpha$ can be expressed in the following form:

$$
\begin{align*}
& p_{1}=\mu^{1 / 3} Y_{1}(\mu) \cos 2 \tau+\mu^{1 / 3} Y_{2}(\mu) \sin 2 \tau-\mu x_{0}{ }^{\prime} / b^{*} B_{1}{ }^{*}+\mu x_{1}{ }^{*} M_{3} \cos \tau+\ldots \\
& q_{1}=\mu^{1 / 2} 2 Y_{1}(\mu)\left(A_{1}^{*}\right)^{-1} \sin 2 \tau-\mu^{1 / 3} 2 Y_{2}(\mu)\left(A_{1}^{*}\right)^{-1} \cos 2 \tau+\mu y_{0}^{\prime} / a^{*} A_{1}^{*}+ \\
& 4 \mu x_{2}{ }^{*}\left(A_{1}{ }^{*}\right)^{-1} \sin \tau+\ldots \\
& r_{1}=1-\mu^{2} M_{3}\left[x_{0}{ }^{\prime}(1-\cos \tau)+y_{0}{ }^{\prime} \sin \tau\right]+. \\
& \gamma_{1}=\left(M_{3}+\beta_{3}\right) \cos \tau+\mu^{\epsilon_{3}} v^{*} Y_{1}(\mu) \cos 2 \tau+\mu^{d_{s}} v^{*} Y_{2}(\mu) \sin 2 \tau+\ldots  \tag{3.10}\\
& \gamma_{1}^{\prime}=-\left(M_{3}+\beta_{8}\right) \sin \tau-\mu^{6 / 3} 2 v_{2} * Y_{1}(\mu) \sin 2 \tau 4 \mu^{4 / 2} 2 v_{2}^{*} Y_{2}(\mu) \cos 2 \tau 4 \ldots \\
& \gamma_{1}{ }^{\prime \prime}=1+\mu^{4 / 2} M_{3}\left[a^{*} Y_{1}(\mu)+1 / 2 Y_{1}(\mu)\left(2 b^{*} / A_{1}{ }^{*}-a^{*}\right) \cos \tau+1 /{ }_{2} Y_{2}(\mu)\left(2 b^{*} / A_{1}{ }^{*}-\right.\right. \\
& \left.\left.-a^{*}\right) \sin \tau-1 / 2 Y_{1}(\mu)\left(2 b^{*} / A_{1}{ }^{*}+a^{*}\right) \cos 3 \tau-1 / 3 Y_{2}(\mu)\left(2 b^{*} / A_{1}{ }^{*}+a^{*}\right) \sin 3 \tau\right]+\ldots \\
& \alpha=\mu^{2} \pi\left\{2 M_{3} x_{0}{ }^{\prime}-2 z_{0}{ }^{\prime}-\mu^{2 / 3}\left[Y_{1}{ }^{2}(\mu)\left(a^{*}+1 / 2 a^{*} B_{1}{ }^{*}+2\left(A_{1}{ }^{*}\right)^{-2}-2 b^{*}\left(A_{1}{ }^{*}\right)^{-2}\right)+\right.\right. \\
& \left.\left.+Y_{2}{ }^{2}(\mu)\left(4 b^{*}\left(A_{1}{ }^{*}\right)^{-2}+1 / 2 a^{*} B_{1}{ }^{*}+2\left(A_{1}^{*}\right)^{-2}-2 b\left(A_{1}^{*}\right)^{-2}\right)\right]\right\}+\ldots \tag{3.11}
\end{align*}
$$

Stars indicate the substitution $a=a^{*}, b=b^{*}$.
4. In order to determine the periodic solutions of system (2.21) when
$\omega=1 / 2$, we obtain from formulas (1.23) and (1.22)

$$
\begin{align*}
& F_{2}{ }^{(0)}= {\left[\left(L(1 / 2)+x_{0}{ }^{\prime} M_{3} L_{2}\right) M_{1}+y_{0}{ }^{\prime} M_{3} L_{3} M_{2}\right] \cos 1 / 2 \tau+} \\
&+\left[\left(L(1 / 2)-x_{0}{ }^{\prime} M_{3} L_{2}\right) M_{2}+y_{0}{ }^{\prime} M_{3} L_{3} M_{1}\right] \sin 1 / 2 \tau+\ldots  \tag{4.1}\\
& \Phi_{2}{ }^{(0)}=M_{3} N(1 / 2) \cos \tau+\ldots, L_{2}=-1 / 2 A_{1}\left(a / b-1 / 2 A_{1}{ }^{-1}\right), L_{3}=1 / 2\left(1-1 / 2 A_{1}{ }^{-1}\right) a^{-1} \tag{4.2}
\end{align*}
$$

From formulas (2.11) and (4.1) we find the expressions

$$
\begin{gather*}
G_{2}\left(T_{0}\right)=-4 \pi\left[\left(L_{1}(1 / 2)-x_{0}{ }^{\prime}\left(M_{3}+\beta_{3}\right) L_{2}\right) \beta_{2}+y_{0}{ }^{\prime}\left(M_{3}+\beta_{3}\right) L_{3} \beta_{1}\right]  \tag{4.3}\\
\dot{G_{2}}\left(T_{0}\right)=2 \pi\left[\left(L_{1}(1 / 2)+x_{0}{ }^{\prime}\left(M_{3}+\beta_{3}\right) L_{2}\right) \beta_{1}+y_{0}{ }^{\prime}\left(M_{3}+\beta_{3}\right) L_{3} \beta_{2}\right] \\
H_{2}\left(T_{0}\right)=0, \quad \dot{H_{2}}\left(T_{0}\right)=2 \pi\left(M_{3}+\beta_{3}\right) N_{1}(1 / 2) \quad\left(T_{0}=4 \pi\right)
\end{gather*}
$$

which, when substituted into the system (2.21), give the equations for $\beta_{1}$ and $\beta_{2}$

$$
\begin{gathered}
\left\{L_{1}(1 / 2)-1 / 4 N_{1}(1 / 2)-x_{0}{ }^{\prime} M_{3} L_{2}\right] \beta_{2}+y_{0}{ }^{\prime} M_{3} L_{3} \beta_{1}+\mu\left[-1 / 4 \pi^{-1} g_{3}\left(T_{0}\right)+\right. \\
\left.+\rho_{1}\left(\beta_{1}, \beta_{2}\right)\right]+\mu^{2}(\ldots)=0 \\
{\left[L_{1}\left({ }^{1} / 2\right)-1 / 4 N_{1}(1 / 2)+x_{0}{ }^{\prime} M_{3} L_{2}\right] \beta_{1}+y_{0}{ }^{\prime} M_{3} L_{3} \beta_{2}+\mu\left[1 / 2 \pi^{-1} \dot{g}_{3}\left(T_{0}\right)+\rho_{2}\left(\beta_{1}, \beta_{2}\right)\right]+} \\
+\mu^{2}(\ldots)=0
\end{gathered}
$$

By formula (2.23), and under conditions $g_{3}(4 \pi)=\dot{g}_{3}(4 \pi)=0$ which are satisfied when $\omega=1 / 2$, system (4.4) can be expressed in the form

$$
\begin{gathered}
y_{0}{ }^{\prime} M_{3} L_{3} \beta_{1}-\left[x_{0}^{\prime} M_{3} L_{2}+z_{0}^{\prime} W_{1}(1 / 2)\right] \beta_{2}+W\left({ }^{1} / 2\right) \beta_{2}\left(\beta_{1}{ }^{2}+\beta_{2}{ }^{2}\right)+\mu \rho_{3}\left(\beta_{1}, \beta_{2}\right)+ \\
+\mu^{2}(\ldots)=0 \\
{\left[x_{0}^{\prime} M_{3} L_{2}-z_{0}^{\prime} W_{1}\left({ }^{1 / 2}\right)\right] \beta_{1}+y_{0}^{\prime} M_{3} L_{3} \beta_{2}+W(1 / 2) \beta_{1}\left(\beta_{1}{ }^{2}+\beta_{2}{ }^{2}\right)+\mu \rho_{4}\left(\beta_{1}, \beta_{2}\right)+} \\
+\mu^{2}(\ldots)=0
\end{gathered}
$$

where $\rho_{i}\left(\beta_{1}, \beta_{2}\right)(i=1,2,3,4)$ are entire functions of their arguments vanishing at $\beta_{1}=\beta_{2}=0$. We shall assume that the following condition:

$$
\begin{equation*}
M_{3}{ }^{2}\left(x_{0}{ }^{\prime 2} L_{2}{ }^{2}+y_{0}{ }^{\prime 2} L_{3}{ }^{2}\right)-z_{0}{ }^{\prime} W_{1}{ }^{2}(1 / 2) \neq=0 \tag{4.6}
\end{equation*}
$$

is satisfied.
By the inequality $a>b>1$, which results from the condition $\omega=1 / 2$, we obtain from formulas (4.2) that $L_{2}>0, L_{3}>0$; besides, $W_{1}(1 / 2) \neq 0$.

Consequently, for any selection of numbers $x_{0}{ }^{\prime}, y_{0}{ }^{\circ}$ and $z_{0}{ }^{\prime}$ satisfying the condition

$$
\begin{equation*}
z_{0}^{\prime}\left(x_{0}^{\prime 2}+y_{0}^{\prime 2}\right) \neq 0 \tag{4.7}
\end{equation*}
$$

and any selection of the quantities $a$ and $b$ satisfying the relation
$\omega=1 / 2$, there exists only one value $M_{3}$

$$
\begin{equation*}
M_{3}^{\circ}=\left[z_{0}^{\prime} W_{1}^{2}(1 / 2) /\left(x_{0}^{\prime 2} L_{2}^{2}+y_{0}^{\prime 2} L_{3}^{2}\right)\right]^{1 / k} \tag{4.8}
\end{equation*}
$$

or, by (2.17) and (1.3), one value $\gamma_{0}{ }^{\prime \prime}$

$$
\begin{equation*}
\Gamma_{0}^{\prime \prime}=\left(x_{0}{ }^{\prime 2} L_{2}^{2}+y_{0}^{\prime 2} L_{3}{ }^{2}\right)^{1 / 2}\left[x_{0}^{\prime 2} L_{2}{ }^{2}+y_{0}^{\prime 2} L_{3}^{2}+z_{0}^{\prime 2} W_{1}^{2}\left(y_{2}\right)\right]^{-1 / 2} \tag{4.9}
\end{equation*}
$$

at which condition (4.6) is not satisfied. If condition (4.7) is not satisfied, relation (4.6) is satisfied for any value of $M_{3}$. In this way condition (4.6) is equivalent to the condition

$$
\begin{equation*}
\tau_{0}{ }^{\prime \prime} \neq \Gamma_{0}{ }^{\prime \prime} \quad\left(\text { or } M_{3} \neq M_{3}{ }^{\circ}\right) \tag{4.10}
\end{equation*}
$$

When condition (4.10) is satisfied, then from equations (4.5) it follows [7] that $\beta_{1}$ and $\beta_{2}$, expanded in power series in $\mu$, begin from a term of order not higher than $\mu^{2}$. In this case, when $\omega=1 / 2$, the first terms of the power series expansion of $\mu$ of the periodic solution of system (1.6) and the $\alpha^{\prime} s$ can be found from formulas (2.30) and (2.31).
5. Let us consider the case $\omega=1$, which corresponds to the disc (1.15). In this case we must set in formulas (1.18) to (1.23)

$$
\begin{equation*}
x_{1}=x_{2}=v=0, \quad v_{2}=B_{1}=-1, \quad A_{1}=1 \tag{5.1}
\end{equation*}
$$

Then, substituting in formulas (1.22) and (1.23) for $S_{i j}^{(0)}(i, j=$ 1. 2)
$S_{11}{ }^{(0)}=-2 M_{3}\left[x_{0}{ }^{\prime}(1-\cos \tau)+y_{0}{ }^{\prime} \sin \tau\right]-E^{2}(a-b)\left[1 / 2 \cos 2(\tau-\varepsilon)+1 / 2-\cos ^{2} \varepsilon\right]$
$S_{12}{ }^{(0)}=E\left\{y_{0}{ }^{\prime}(1-b / a)[\sin (\tau-\varepsilon)+\sin \varepsilon]-a b^{-1} x_{0}{ }^{\prime}[\cos (\tau-\varepsilon)-\cos \varepsilon]\right\}$
$S_{21}{ }^{(0)}=1 / 2 M_{3} E[-\cos (2 \tau-\varepsilon)+\cos \varepsilon]$
$S_{22}{ }^{(0)}=M_{3}\left(-a b^{-1} x_{0}{ }^{\prime} \cos \tau+b a^{-1} y_{e^{\prime}} \sin \tau+a b^{-1} x_{0}{ }^{\prime}\right)+E^{2} b\left[1 / 2 \cos 2(\tau-\varepsilon)-1 / 2+\sin ^{2} \varepsilon\right]$ and also by (2.11) and (2.20) we obtain

$$
\begin{gather*}
G_{2}\left(T_{0}\right)=-\pi \beta_{2}\left[2\left(M_{3}+\beta_{3}\right) x_{0}{ }^{\prime}+1 / 2(a-b)\left(\beta_{2}{ }^{2}-\beta_{1}{ }^{2}\right)\right]  \tag{5.3}\\
\dot{G}_{2}\left(T_{0}\right)=\pi \beta_{1}\left[2\left(M_{3}+\beta_{3}\right) x_{0}{ }^{\prime}+1 / 2(a-b)\left(\beta_{\beta^{2}}-\beta_{1}{ }^{2}\right)\right] \\
G_{3}\left(T_{3}\right)=-\pi\left\{-2 \beta_{2}\left[y_{0}^{\prime}\left(1-a^{-1} b\right) \beta_{2}+a b^{-1} x_{0}^{\prime} \beta_{1}\right]+3 a^{-1} b^{-1} x_{0}^{\prime} y_{0}^{\prime}\left(M_{3}+\beta_{3}\right)\right\} \\
\dot{G}_{3}\left(T_{0}\right)=\pi\left\{-2 \beta_{1}\left[y_{0}^{\prime}\left(1-b a^{-1}\right) \beta_{2}+a b^{-1} x_{0}^{\prime} \beta_{1}\right]+\left[a^{-2}(a+1) y_{0}^{\prime 2}-\right.\right. \\
\\
\left.\left.-b^{-2}(b+1) x_{0}^{\prime 2}\right]\left(M_{3}+\beta_{3}\right)\right\} \\
H_{2}\left(T_{0}\right)=0, \quad \dot{H}_{2}\left(T_{0}\right)=\pi\left(M_{3}+\beta_{3}\right)\left[2\left(M_{3}+\beta_{3}\right) x_{0}^{\prime}-\left(a \beta_{1}{ }^{2}+b 3_{2}{ }^{2}\right)\right] \quad\left(T_{0}=2 \pi\right)
\end{gather*}
$$

Substituting formulas (5.3) in system (2.21) and in (2.19) when $\omega=1$, we obtain the equations determining $\beta_{1}$ and $\beta_{2}$ and the expression for $\alpha$ in the form

$$
\begin{gather*}
\beta_{1}\left(\beta_{1}^{2}+\beta_{2}^{2}\right)=-\mu V_{3}+\ldots, \beta_{2}\left(\beta_{1}^{2}+\beta_{2}{ }^{2}\right)=-\mu V_{4}+\ldots  \tag{5.4}\\
V_{3}=2 M_{3}\left[a^{-2}(a+1) y_{0}^{\prime 2}-b^{-2}(b+1) x_{0}{ }^{\prime 2}\right], \quad V_{4}=6 M_{3} a^{-1} b^{-1} x_{0}^{\prime} y_{0}^{\prime} \tag{5.5}
\end{gather*}
$$

$$
\begin{equation*}
\alpha=\mu^{2} \pi\left(2 M_{3} x_{0}^{\prime}-a \beta_{1}^{2}-b \beta_{2}^{2}+\ldots\right) \tag{5.6}
\end{equation*}
$$

Whence

$$
\begin{array}{ll}
\beta_{1}=\mu^{1 / 3} Y_{3}(\mu), & Y_{3}(\mu)=-V_{3}\left[V_{3}^{2}+V_{4}^{2}\right]^{-1 / 3}+\mu^{1 / 3}(\ldots)  \tag{5.7}\\
\beta_{2}=\mu^{1 / 3} Y_{4}(\mu), & Y_{4}(\mu)=-V_{4}\left[V_{3}^{2}+V_{4}^{2}\right]^{-1 / 3}+\mu^{1 / 3}(\ldots)
\end{array}
$$

From formulas (5.5) it follows that the quantities $V_{3}$ and $V_{4}$ do not vanish simultaneously; therefore the relations (5.7) are always true, with the exception of the case $x_{0}^{\prime}=y_{0}^{\prime}=z_{0}^{\prime}=0$, which we have not considered. Consequently, when $\omega=1$, the first terms of the expansion of the periodic solution of system (1.6) and the $a^{*} s$ can be expressed in the form

$$
\begin{gather*}
p_{1}=\mu^{1 / 3} Y_{3}(\mu) \cos \tau+\mu^{1 / 3} Y_{4}(\mu) \sin \tau+\mu x_{0}^{\prime} / b+\ldots \\
q_{1}=\mu^{1 / 3} Y_{3}(\mu) \sin \tau-\mu^{1 / 3} Y_{4}^{\prime}(\mu) \cos \tau+\mu y_{0}^{\prime} / a+\ldots  \tag{5.8}\\
r_{1}=1-\mu^{2} M_{3}\left[x_{0}^{\prime}(1-\cos \tau)+y_{0}^{\prime} \sin \tau\right]+\ldots, \gamma_{1}=\left(M_{3}+\beta_{3}\right) \cos \tau+\ldots \\
\gamma_{1}^{\prime}=-\left(M_{3}+\beta_{3}\right) \sin \tau+\mu^{1 / 3} Y_{3}(\mu) \sin \tau-\mu^{4 / 3} Y_{4}(\mu) \cos \tau+\ldots \\
\gamma_{1}^{\prime \prime}=1-{ }^{1 / 2} \mu^{4 / 3} M_{3}\left[Y_{3}(\mu) \cos 2 \tau+Y_{4}(\mu) \sin 2 \tau-Y_{3}(\mu)\right]+\ldots \\
\alpha=\mu^{2} \pi\left[2 M_{3} x_{8}^{\prime}-\mu^{2 / 3} a Y_{3}^{2}(\mu)-\mu^{2 / 3} b Y_{4}^{2}(\mu)+\ldots\right] \tag{5.9}
\end{gather*}
$$

6. In the case when $\omega$ is irrational, on the strength of conditions (2.17) and (2.20), we must $[8]$ search for a periodic solution of system (1.21) whose period is $2 \pi+\alpha$, in the form

$$
\begin{gather*}
p_{2}(\tau, \mu)=\chi_{1}\left(\beta_{3}, \mu\right) \cos \omega \tau+\chi_{2}\left(\beta_{3}, \mu\right) \sin \omega \tau+\sum_{k=2}^{\infty} G_{k}(\tau) \mu^{k}  \tag{6.1}\\
\gamma_{2}(\tau, \mu)=\left(M_{3}+\beta_{3}\right) \cos \tau+\sum_{k=2}^{\infty} H_{k}(\tau) \mu^{k}
\end{gather*}
$$

where $X_{1}\left(\beta_{3}, \mu\right)$ and $X_{2}\left(\beta_{3}, \mu\right)$ are analytic functions to be determined; besides. $X_{1}\left(\beta_{3}, 0\right)=X_{2}\left(\beta_{3}, 0\right)=0$. Writing these functions as the series

$$
\chi_{i}\left(\beta_{3}, \mu\right)=\sum_{k=1}^{\infty}\left(Q_{k}^{(i)}+\frac{\partial Q_{k}^{(i)}}{\partial M_{3}} \beta_{3}+1 / 2 \frac{\partial^{2} Q_{k}^{(i)}}{\partial M_{3^{2}}^{2}} \beta_{3}^{2}+\ldots\right) \mu^{k} \quad(i=1,2)(6.2)
$$

and replacing $M_{1}+\beta_{1}$ and $M_{2}+\beta_{2}$ by the above functions under the conditions of periodicity (2.18), we obtain by (6.1) an infinite system of equations determining the coefficients $Q_{k}{ }^{(i)}$. Since the coefficients $Q_{1}{ }^{(i)}$ are determined from the system of linear homogeneous algebraic equations
$Q_{1}^{(1)}(1-\cos 2 \pi \omega)-Q_{1}^{(2)} \sin 2 \pi \omega=0, \quad Q_{1}^{(1)} \sin 2 \pi \omega+Q_{1}^{(2)}(1-\cos 2 \pi \omega)=0$
with non-zero determinant $\Delta=2(1-\cos 2 \pi \omega)$, therefore they are trivial solutions of the above system, and series (6.2) begins from a term of order not lower than $\mu^{2}$. It follows that the expansion of $p_{2}(T, \mu)$ in power series in $\mu$ also begins from a term of order not lower than $\mu^{2}$; the quantities $H_{2}\left(T_{0}\right)$ and $\dot{H}_{2}\left(T_{0}\right) \quad\left(T_{0}=2 \pi\right)$ can be found from formula (2.12) by replacing $N(\omega)$ by $N_{1}(\omega)$ and $M_{3}$ by $M_{3}+\beta_{3}$.

In this way the first terms of the expansion in power series of $\mu$ of the periodic solution of system (1.6) and the quantity $\alpha$ can be found from formulas (2.30) and (2.31).
7. We shall use now the Eulerian angles $\theta, \psi$ and $\varphi$ when analyzing the obtained motion of a heavy solid about a fixed point

$$
\begin{equation*}
\cos \theta=\gamma^{\prime \prime}, \quad \frac{d \psi}{d t}=\frac{p \gamma+q \gamma^{\prime}}{1-\gamma^{\prime 2}}, \quad \frac{d \varphi}{d t}=r-\frac{d \psi}{d t} \cos \theta, \quad \tan \varphi_{0}=\frac{\gamma_{0}}{\gamma_{0}^{\prime}} \tag{7.1}
\end{equation*}
$$

Let us mention in advance that, since the initial system (1.1) and system (1.21) are autonomous, the periodic solutions will remain periodic if $t$ is replaced by $t+h$, where $h$ is arbitrary. By formula (7.1), and by taking into account that the initial instant of time corresponds to the instant $t=h$, we can replace $h$ by

$$
\begin{equation*}
\varphi_{0}=1 / 2 \pi+r_{0} h+\cdots \tag{7.2}
\end{equation*}
$$

introducing in this way an arbitrary initial angle of spin.
We have obtained previously four sets of formulas, (2.30), (3.7), (3.10) and (5.8), which exhaust all forms of different expansions of the periodic solutions with the prescribed approximations and restrictions. Substituting in (7.1) expansions (2.30), (3.7). (3.10) and (5.3), in which $t$ has been replaced by $t+h$, and relations (1.5) and by using (2.17) in the form

$$
\begin{equation*}
M_{3}=\tan \theta_{0} \tag{7.3}
\end{equation*}
$$

we obtain the following expressions for the angles $\theta$ and $\psi$. In the set (2.30)

$$
\begin{equation*}
\theta-\theta_{0}=-\mu^{2}\left[\theta^{(1)}(t+h)-\theta^{(1)}(h)\right]+\ldots \tag{7.4}
\end{equation*}
$$

$$
r_{0}\left(\psi-\psi_{0}\right)=-\mathrm{Mg}^{-1} z_{0} t+\mu c \operatorname{cosec} \theta_{0} \sqrt{\cos \theta_{0}}\left[\psi^{(1)}(t+h)-\psi^{(1)}(h)\right]+\ldots
$$

$$
\theta^{(1)}(t)=y_{0}^{\prime} a^{-1} A_{1}^{-1} \sin r_{0} t+x_{0}^{\prime} b^{-1} B_{1}^{-1} \cos r_{0} t-1 / 2 z_{0}^{\prime} \tan \theta_{0}\left(A_{1}-1\right)\left(A_{1}+1\right)^{-1} \cos 2 r_{0} t
$$

$$
\psi^{(1)}(t)=-x_{0}^{\prime} b^{-1} B_{1}^{-1} \sin r_{0} t+y_{0}^{\prime} a^{-1} A_{1}^{-1} \cos r_{0} t+1 / 4 \tan \theta_{0}\left(x_{1}+x_{2} / A_{1}\right) \sin 9 r_{0} t
$$

In the set (3.7)

$$
\begin{equation*}
\theta-\theta_{0}=-\mu^{2}\left[\theta^{(2)}(t \psi h)-\theta^{(2)}(h)\right] 4 \ldots \tag{7.5}
\end{equation*}
$$

$r_{0}\left(\psi-\psi_{0}\right)=-\operatorname{Mg} C^{-1} z_{0} t+\mu c \operatorname{casec} \theta_{0} \sqrt{\cos \theta_{0}}\left[\psi^{(2)}(t+h)-\psi^{(2)}(h)\right] 4 \ldots$

$$
\begin{gathered}
\theta^{(2)}(t)=s_{1} \sin r_{0} t+s_{2} \cos r_{0} t+s_{3} \cos 2 r_{0} t+s_{4} \cos 3 r_{0} t+s_{5} \sin 3 r_{0} t \\
\psi^{(2)}(t)= \\
=-\left[\frac{x_{0}^{\prime}}{b B_{1}}-\left(\frac{1}{2}-\frac{1}{A_{1}}\right) \frac{V_{1}}{z_{0}^{\prime} W_{1}(2)}\right] \sin r_{0} t+\left[\frac{y_{0}^{\prime}}{a A_{1}}-\left(\frac{1}{2}-\frac{1}{A_{1}}\right) \frac{V_{2}}{z_{0}^{\prime} W_{1}(2)}\right] \cos r_{0} t+ \\
+\frac{\tan \theta_{0}}{4}\left(x_{1}+\frac{x_{2}}{A_{1}}\right) \sin 2 r_{0} t+\frac{1}{3}\left(\frac{1}{2}+\frac{1}{A_{1}}\right) \frac{V \sin 3 r_{0} t}{z_{0}^{\prime} W_{1}(2)}-\frac{1}{3}\left(\frac{1}{2}+\frac{1}{A_{1}}\right) \frac{V_{2} \cos 3 r_{0} t}{z_{0}{ }^{\prime} W_{1}(2)}
\end{gathered}
$$

In the set (3.10)

$$
\begin{gathered}
\theta-\theta_{0}=-\mu^{4 / 3}\left[\theta^{(3)}(t+h)-\theta^{(3)}(h)\right]+\ldots \\
r_{0}\left(\psi-\psi_{0}\right)=\left\{-M g z_{0} C^{-1}+1 / 2 \mu^{2 / 3}\left[Y_{\mathbf{1}}{ }^{2}(\mu)+Y_{2}{ }^{2}(\mu)\right]\left(v^{*}-4 v_{2}{ }^{*} / A_{1}{ }^{*}\right)\right\} t+ \\
4 \mu^{1 / 3} c \operatorname{cosec} \theta_{0} \sqrt{\cos \theta_{0}}\left[\psi^{(3)}(t+h)-\psi^{(3)}(h)\right]+\ldots
\end{gathered}
$$

$$
\theta^{(3)}(t)=1 / 2 Y_{1}(0)\left(2 b^{*} / A_{1}^{*}-a^{*}\right) \cos r_{0} t+1 / 2 Y_{2}(0)\left(2 b^{*} / A_{1}^{*}-a^{*}\right) \sin r_{0} t-
$$

$$
-1 / 2 Y_{1}(0)\left(2 b^{*} / A_{1}^{*}+a^{*}\right) \cos 3 r_{0} t-1 / 2 Y_{2}(0)\left(2 b^{*} / A_{1}{ }^{*}+a^{*}\right) \sin 3 r_{0} t
$$

$$
\psi^{(3)}(t)=\left(1 / 2-1 / A_{1}^{*}\right) Y_{1}(0) \sin r_{0} t-\left(1 / 2-1 / A_{1}{ }^{*}\right) Y_{2}(0) \cos r_{0} t+
$$

$$
+1 / 3\left(1 / 2+1 / A_{1}{ }^{*}\right) Y_{1}(0) \sin 3 r_{0} t-1 / 3\left(1 / 2 \neq 1 / A_{1}{ }^{*}\right) Y_{2}(0) \cos 3 r_{0} t
$$

In the set (5.8)

$$
\begin{gathered}
\theta-\theta_{0}=-1 / 2 \mu^{1 / 3}\left[\theta^{(4)}(t+h)-\theta^{(4)}(h)\right]+\ldots \\
r_{0}\left(\psi-\psi_{0}\right)=1 / 2 \mu^{1 / 3} c \operatorname{cosec} \theta_{0} \sqrt{\cos \theta_{0}}\left[\psi^{(4)}(t+h)-\psi^{(4)}(h)\right]+\ldots
\end{gathered}
$$

$\theta^{(4)}(t)=-Y_{3}(0) \cos 2 r_{0} t-Y_{4}(0) \sin 2 r_{0} t-\psi^{(4)}(t)=Y_{3}(0) \sin 2 r_{0} t-Y_{4}(0) \cos 2 r_{0} t$
In all the cases $\varphi$ is expressed by the formula

$$
\begin{equation*}
\varphi-\varphi_{0}=\left[r_{0}+M g C^{-1} r_{0}{ }^{-1}\left(z_{0} \cos \theta_{0}-x_{0} \sin \theta_{0}\right)\right] t+\ldots \tag{7.8}
\end{equation*}
$$

In all these formulas $\theta_{0}$, as indicated by formula (7.3), is an arbitrary initial (at $t=0$ ) angle of nutation; $\psi_{0}$ is an arbitrary initial angle of precession. Replacing $h$ by $\varphi_{0}$ we can see from (7.2) that the expressions for the Eulerian angles $\theta, \psi$ and $\varphi$ depend on four arbitrary constants $\theta_{0} . \psi_{0} . \varphi_{0}$ and $r_{0}\left(r_{0}\right.$ is large).

From formulas (7.4) to (7.6) it follows that for a heavy solid, when $z_{0} \neq 0$ (with the exception of the special cases discussed), with one point fixed, which spins rapidly about the major or the minor axis of the ellipsoid of inertia, it is possible to show initial conditions at which the solid will perform, in the first approximation, a pseudoregular precession about the vertical axis. Four out of these six initia conditions, $\theta_{0}, \Psi_{0}, \varphi_{0}$ and $r_{0}\left(r_{0}\right.$ is large), can be arbitrary.

As an example, we shall consider the case of a regular precession of a rapidly spinning Lagrange gyroscope ( $A=B, x_{0}=y_{0}=0$ ). System (1.1) rewritten for this case, will have four particular solutions

$$
\begin{gather*}
\gamma^{\prime \prime}=\gamma_{0}^{\prime \prime}, \quad p=\lambda \Upsilon, \quad q=\lambda \gamma^{\prime}, \quad r=r_{0}  \tag{7.9}\\
\lambda=\frac{r_{0}\left(m_{1}-1\right)}{2 \gamma_{0}^{\prime \prime}}\left[-1+\left(1+\frac{4 \gamma_{0}^{\prime \prime} c_{1}{ }^{2}}{r_{0}^{2}\left(m_{1}-1\right)^{2}}\right)^{1 / 2}\right], \quad m_{1}=\frac{A-C}{A}, \quad c_{1}{ }^{2}=\frac{M g z_{0}}{A}
\end{gather*}
$$

Substituting solutions (7.9) into (7.1), We obtain the relations

$$
\cos \theta=\gamma_{0}^{\prime \prime}, \quad d \Psi / d t=\lambda, \quad d \varphi / d t=r_{0}-\lambda \gamma_{0}^{\prime \prime}
$$

for which, taking into account that $r_{0}$ is large, we obtain
$\theta=\theta_{0}, \quad \psi=\psi_{0}-\mathrm{Mg} C^{-1} r_{0}^{-1} z_{0} t+\ldots, \varphi=\varphi_{0}+\left(r_{0}+\mathrm{Mg}^{-1} r_{0}^{-1} z_{0} \cos \theta_{0}\right) t+\ldots$
Expressions (7.10) coincide with formulas (7.4), (7.5), (7.6) and (7.8) when they are rewritten for the considered case.

Formulas (7.7) are for the case $\omega=1$, under the condition that $z_{0}=0$. Under this last condition the restrictions caused by retaining only the first terms in the expansion of the required periodic solution permit also an investigation of the cases of irrational $\omega$ and of $\omega=1 / 2$. Formula (7.4) takes care of these two cases. Formulas (7.4) and (7.7) will be rewritten now in the following form:

$$
\begin{gathered}
\theta-\theta_{j}=R_{j} \sin \theta_{0} \cos \left(j r_{0} t-\varepsilon_{j}\right)+\ldots \quad(j=1,2), \quad R_{1}=\mu^{2} E_{1} \operatorname{cosec} \theta_{0} \\
\psi-\psi_{j}=R_{j} \sin \left(j r_{0} t-\varepsilon_{j}\right)+\ldots, \quad R_{2}=1 / \mu^{4 / 3} E_{2} \operatorname{cosec} \theta_{0} \\
\varphi-\varphi_{0}=\left(r_{0}-M g C^{-1} r_{0}^{-1} x_{0} \sin \theta_{0}\right) t+\ldots \\
E_{1}=\left[\left(\frac{x_{0}^{\prime}}{a-1}\right)^{2}+\left(\frac{y_{0}^{\prime}}{b-1}\right)^{2}\right]^{1 / 2}, \quad E_{2}=\left[Y_{3}^{2}(0)+Y_{4}^{2}(0)\right]^{1 / 2}, \quad \varepsilon_{j}=\varepsilon_{j}^{\circ}-j\left(\varphi_{0}-1 / 2 \pi\right) \\
\cos \varepsilon_{1}^{\circ}=\frac{x_{0}^{\prime}}{(1-a) E_{1}}, \quad \sin \varepsilon_{1}^{\circ}=\frac{y_{0}^{\prime}}{(b-1) E_{1}}, \quad \cos \varepsilon_{2}^{\circ}=\frac{Y_{3}(0)}{E_{2}} \quad \sin \varepsilon_{2}^{\circ}=\frac{Y_{4}(0)}{E_{2}} \\
\theta_{j}=\theta_{0}-R_{j} \sin \theta_{0} \cos \varepsilon_{j}, \quad \psi_{j}=\psi_{0}+R_{j} \sin \varepsilon_{j}
\end{gathered}
$$

Where $j=2$ is for the case of a disc.
Let us consider a spherical rectangle bounded by parallels distant from the middle parallel $\theta_{j}$ by the angle $\pm R_{j} \sin \theta_{0}$. and by meridians distant from the middle meridian $\psi_{j}$ by the angle $\pm R_{j}$. Then the trajectory of the $z$-axis of our solid will be the ellipse

$$
\begin{equation*}
\frac{\left(\theta-\theta_{j}\right)^{2}}{R_{j}^{2} \sin ^{2} \theta_{0}}+\frac{\left(\psi-\psi_{j}\right)^{2}}{R_{j}^{2}}=1 \quad(j=1,2) \tag{7.12}
\end{equation*}
$$

tangent to the sides of this spherical rectangle at their midpoints.
The z-axis of the body tracing this ellipse performs, in the first approximation, a periodic motion whose period is $T=2 \pi / j r_{0}$ and at the instants of time

$$
t_{n_{1}}=\left(\pi n_{1}+\varepsilon_{j}\right) / j r_{0} \quad\left(n_{1}=0, \pm 1, \pm 2, \ldots\right)
$$

it will pass through the points of intersection of the middle parallel and the bounding meridians, and at the instants

$$
t_{n_{2}}=\left[1 / 2\left(2 n_{2}+1\right) \pi+\varepsilon_{j}\right] / j r_{0} \quad\left(n_{2}=0, \pm 1, \pm 2, \ldots\right)
$$

it will pass through the points of intersection of the middle meridian with the bounding parallels. Formula (7.11) shows that the spin differs very little from the uniform rotation with large angular velocity $r_{0}$.

Thus, for a heavy solid with one point fixed, with $z_{0}=0$, with the described restrictions on the moments of inertia, which spins rapidly about the major or the minor axis of the ellipsoid of inertia, we can show initial conditions at which the body, in the first approximation, will perform the motion which we have just found. In this case, as in the case when $z_{0} \neq 0$, out of the six initial conditions, at least four, $\theta_{0}, \psi_{0}, \varphi_{0}$ and $r_{0}\left(r_{0}\right.$ is large) can be arbitrary.

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